

Neutrino Oscillation in Matter

We discuss next the topics of "Neutrino Oscillation in Matter"

Consider $H = H_0 + W$, where

$$H_0 |\phi_1\rangle = m_1^0 |\phi_1\rangle, \quad H_0 |\phi_2\rangle = m_2^0 |\phi_2\rangle$$

$|\phi_1\rangle, |\phi_2\rangle$ are the eigenstates of H_0

W is represented by a Hermitian matrix in the basis of $|\phi_1\rangle, |\phi_2\rangle$

$$W = \begin{pmatrix} W_{11} & W_{12} \\ W_{21} & W_{22} \end{pmatrix} \quad W_{12} = W_{21}^*$$

and H in the basis of $|\phi_1\rangle, |\phi_2\rangle$ becomes

$$H = \begin{pmatrix} m_1' & m_{\text{I}}^* \\ m_{\text{I}} & m_2' \end{pmatrix}$$

The eigenstates of H , $|\psi_1\rangle, |\psi_2\rangle$ can be expressed in terms of $|\phi_1\rangle, |\phi_2\rangle$ as

$$|\psi_1\rangle = \cos\theta |\phi_1\rangle + \sin\theta |\phi_2\rangle$$

$$|\psi_2\rangle = -\sin\theta |\phi_1\rangle + \cos\theta |\phi_2\rangle$$

H can be diagonalized as follows

$$\begin{pmatrix} \cos\theta & -\sin\theta \\ \sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} m_1' & m_{\text{I}}^* \\ m_{\text{I}} & m_2' \end{pmatrix} \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} = \begin{pmatrix} m_1 & 0 \\ 0 & m_2 \end{pmatrix}$$

One can obtain easily

$$\tan 2\theta = \frac{2|m_{\pm}|}{m_2' - m_1'}$$

$$m_2 = \frac{1}{2} \left[m_1' + m_2' + \sqrt{(m_2' - m_1')^2 + 4m_{\pm}^2} \right]$$

$$m_1 = \frac{1}{2} \left[m_1' + m_2' - \sqrt{(m_2' - m_1')^2 + 4m_{\pm}^2} \right]$$

We note that

$$m_1 + m_2 = m_1' + m_2'$$

$$m_2 - m_1 = \sqrt{(m_2' - m_1')^2 + 4m_{\pm}^2} > (m_2' - m_1')$$

θ can reach $\pi/4$ if w_{11} , w_{22} are adjusted such that

$$m_2' - m_1' = m_2^0 + w_{22} - m_1^0 - w_{11} = 0$$

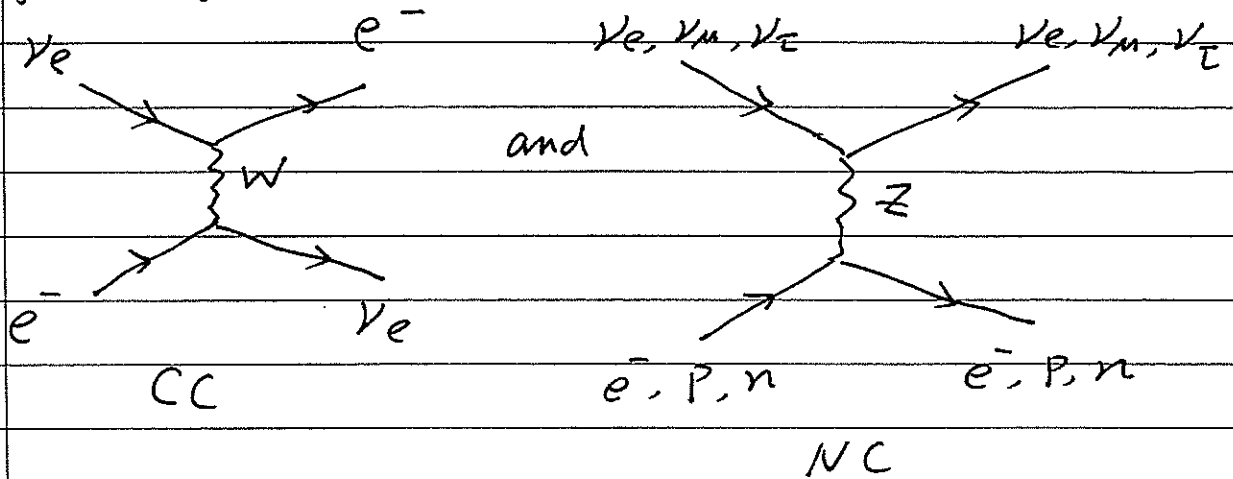
$$\text{or } \frac{|w_{12}|}{m_2' - m_1'} \rightarrow \infty$$

For neutrino in vacuum, then $w = 0$, $H = H_0$
 $|\phi_1\rangle$, $|\phi_2\rangle$ are the mass eigenstates.

For neutrino in matter, then $w \neq 0$

and the angle θ is non-zero for $|w_{12}| > 0$

neutrinos can undergo coherent forward elastic scatterings via CC or NC via the following Feynman diagrams:



At low energies, the W/Z propagators become constant, and the 'point' interaction gives

$$\mathcal{L}^{CC} = -\frac{G_F}{\sqrt{2}} \bar{\psi}_\mu \gamma^\mu \psi^+$$

or, the effective CC Hamiltonian is

$$H^{CC}(x) = \frac{G_F}{\sqrt{2}} [\bar{\nu}_e(x) \gamma^\mu (1-\gamma^5) e(x)] [\bar{e}(x) \gamma_\mu (1-\gamma^5) \nu_e(x)]$$

It would be convenient to interchange $e(x)$ and $\nu_e(x)$, as indicated above.

This can be accomplished using the Fierz transformation

$$\begin{aligned} (\bar{\psi}_1 \Gamma^a \psi_2) (\bar{\psi}_3 \Gamma^a \psi_4) &= -\frac{1}{16} \sum_b \text{Tr}[(\Gamma^a \Gamma^b)^2] (\bar{\psi}_1 \Gamma^b \psi_4) (\bar{\psi}_3 \Gamma^b \psi_2) \\ &= \sum_b C_{ab} (\bar{\psi}_1 \Gamma^b \psi_4) (\bar{\psi}_3 \Gamma^b \psi_2) \end{aligned}$$

From the table of C_{ab} , one can obtain

$$\begin{aligned} & (\bar{\psi}_1 \gamma^\mu (1-\gamma^5) \psi_2) (\bar{\psi}_3 \gamma_\mu (1-\gamma^5) \psi_4) \\ &= (\bar{\psi}_1 \gamma^\mu (1-\gamma^5) \psi_4) (\bar{\psi}_3 \gamma_\mu (1-\gamma^5) \psi_2) \end{aligned}$$

using the Fierz transformation, we obtain

$$\mathcal{H}^{cc}(x) = \frac{G_F}{\sqrt{2}} [\bar{\nu}_e(x) \gamma^\mu (1-\gamma^5) \nu_e(x)] [\bar{e}(x) \gamma_\mu (1-\gamma^5) e(x)]$$

For an unpolarized electron environment with an energy distribution $f(E_e)$, the averaged Hamiltonian becomes

$$\overline{\mathcal{H}^{cc}(x)} = \int d^3 p_e f(E_e) \langle \mathcal{H}^{cc}(x) \rangle$$

where $\langle \mathcal{H}^{cc}(x) \rangle$ is the expectation value of $\mathcal{H}^{cc}(x)$ summed over the helicity $h_e = \pm 1$:

$$\begin{aligned} \langle \mathcal{H}^{cc}(x) \rangle &= \frac{G_F}{\sqrt{2}} [\bar{\nu}_e(x) \gamma^\mu (1-\gamma^5) \nu_e(x)] \\ &\quad \times \frac{1}{2} \sum_{h_e = \pm 1} \langle \bar{e}(p_e, h_e) | \bar{e}(x) \gamma_\mu (1-\gamma^5) e(x) | e(p_e, h_e) \rangle \end{aligned}$$

Note that $e(x)$, $\nu_e(x)$, etc. are operators and single electron state has the normalization

$$|e^-(p_e, h_e)\rangle = \frac{1}{\sqrt{2E_e V}} a_e^\dagger(p_e, h_e) |0\rangle$$

$\int d^3 p f(E_e) = N_e V$, N_e is the electron density, V is the volume, $N_e V$ is the total number of electrons

$$\text{From } e(x) = \int \frac{d^3 p}{(2\pi)^3 2E} \sum_{he=\pm 1} \left[a^{(he)}(p_e) u^{(he)}(p_e) e^{-i p_e \cdot x} + b^{(he)\dagger}(p_e) v^{(he)}(p_e) e^{-i p_e \cdot x} \right]$$

one can obtain

$$\begin{aligned} & \frac{1}{2} \sum_{he=\pm 1} \langle e(p_e, he) | \bar{e}(x) \gamma_\mu (1-\gamma^5) e(x) | e(p_e, he) \rangle \\ &= \frac{1}{4E_e V} \sum_{he=\pm 1} \overline{u_e^{(he)}(p_e)} \gamma_\mu (1-\gamma^5) u_e^{(he)}(p_e) \end{aligned}$$

Note that $\overline{u_e^{(he)}(p_e)}$ and $u_e^{(he)}(p_e)$ have the same value of p_e , as required by coherent forward elastic scattering.

$$\begin{aligned} & \sum_{he=\pm 1} \overline{u_e^{(he)}(p_e)} \gamma_\mu (1-\gamma^5) u_e^{(he)}(p_e) \\ &= \text{Tr} \left[\left(\sum_{he=\pm 1} u_e^{(he)}(p_e) \overline{u_e^{(he)}(p_e)} \right) \gamma_\mu (1-\gamma^5) \right] \\ &= \text{Tr} \left[(\not{p}_e + m_e) \gamma_\mu (1-\gamma^5) \right] \\ &= \text{Tr} \left[\not{p}_e \gamma_\mu + m_e \gamma_\mu - \not{p}_e \gamma_\mu \gamma^5 - m_e \gamma_\mu \gamma^5 \right] \\ &= \text{Tr} \left[\not{p}_e \gamma_\mu \right] = 4 p_{e,\mu} \end{aligned}$$

(Note that $\text{Tr}[\gamma_\mu] = 0$, $\text{Tr}[\gamma_\mu \gamma^5] = 0$, $\text{Tr}[\gamma_\nu \gamma_\mu \gamma^5] = 0$)

$$\overline{H}^{\text{cc}}(x) = \frac{GF}{\sqrt{2}} \frac{1}{V} \int d^3 p_e f(E_e) \overline{\nu_e}(x) \frac{\not{p}_e}{E_e} (1-\gamma^5) \nu_e(x)$$

$$\text{but } \int d^3 p_e f(E_e) \frac{\not{p}_e}{E_e} = \int d^3 p_e f(E_e) \left(\gamma^0 - \frac{\vec{p}_e \cdot \vec{\gamma}}{E_e} \right) = N_e V \gamma^0$$

(since the integrand is odd under $\vec{p}_e \rightarrow -\vec{p}_e$)

Hence,

$$\overline{H}^{CC}(x) = V_{CC} \overline{V_{eL}(x)} \gamma^0 V_{eL}(x)$$

Where $V_{CC} = \sqrt{2} G_F N_e$

Now, we consider the neutral current Hamiltonian,

$$H^{NC}(x) = \frac{G_F}{\sqrt{2}} \sum_{\alpha=e,\mu,\tau} [\overline{V_{\alpha}(x)} \gamma^{\mu} (1-\gamma^5) V_{\alpha}(x)] \sum_f [\overline{f}(x) \gamma_{\mu} (g_V^f - g_A^f \gamma^5) f(x)]$$

f denotes the fermion, with vector and axial vector coupling to Z as g_V^f and g_A^f

$$g_V^f = I_3^f - 2 Q_f \sin^2 \theta_W$$

$$g_A^f = I_3^f$$

f	g_V	g_A
ν	$\frac{1}{2}$	$\frac{1}{2}$

$$e^- \quad -\frac{1}{2} + 2 \sin^2 \theta_W \quad -\frac{1}{2}$$

$$g_{V,A}(p) = -g_{V,A}(e^-)$$

$$u \quad \frac{1}{2} - \frac{4}{3} \sin^2 \theta_W \quad \frac{1}{2}$$

$$g_{V,A}(n) = -g_{V,A}(u)$$

$$d \quad -\frac{1}{2} + \frac{2}{3} \sin^2 \theta_W \quad -\frac{1}{2}$$

$$p(cud) \quad \frac{1}{2} - 2 \sin^2 \theta_W \quad \frac{1}{2}$$

$$n(cudd) \quad -\frac{1}{2} \quad -\frac{1}{2}$$

$$H(p+\bar{e}) \quad 0 \quad 0$$

From the calculation of $\overline{H^{cc}}(x) = V_{cc} \overline{\nu_{eL}^\alpha} \gamma^0 \nu_{eL}(x)$,
 it is clear that the $g_A^f \gamma^5$ term in
 $\sum_f \left[\overline{f}(x) \gamma_\mu (g_V^f - g_A^f \gamma^5) f(x) \right]$ can not contribute.

Therefore, we only need to consider the g_V^f term.

* Unlike the CC Hamiltonian, which is only relevant for ν_e and $f=e$, the NC Hamiltonian applies to ν_e, ν_μ, ν_τ and $f=e, p, n$.

* Since NC does not distinguish ν_e, ν_μ, ν_τ , we only need to label f .

$$V_{cc} = \sqrt{2} G_F N_e \longrightarrow V_{NC}^f = \sqrt{2} G_F g_V^f N_f$$

(note that for V_{cc} , $g_V^f = 1$ for $V-A$ interaction)

* Since the net g_V for $(e^- + p)$ is zero, we only need to consider the contribution from neutrons (which is a small fraction in the SUN, largely consisting of $e^- + p$ plasma)

* Therefore

$$\sum_f V_{NC}^f = -\frac{1}{2} \sqrt{2} G_F N_n$$

is the neutral-current potential for any neutrino flavor α .

Summarizing, the total potentials from NC and CC for ν_e, ν_μ, ν_τ are

$$V(\nu_e) = \sqrt{2} G_F (N_e - \frac{1}{2} N_n)$$

$$V(\nu_\mu) = \sqrt{2} G_F (-\frac{1}{2} N_n)$$

$$V(\nu_\tau) = \sqrt{2} G_F (-\frac{1}{2} N_n)$$

What would be the potential energy of a ν_α (with h helicity) propagating through a medium? It would be the expectation value given as

$$V_\alpha^{(h)} = \langle \nu_\alpha(p, h) | \int d^3x \bar{H}(x) | \nu_\alpha(p, h) \rangle$$

where $\bar{H}(x) = \bar{H}^{CC}(x) + \bar{H}^{NC}(x)$, h is the helicity of ν with flavor α .

$$V_\alpha^{(h)} = \frac{V_\alpha}{4E} \bar{u}_{\nu_\alpha}^{(h)}(p) \gamma^0 (1 - \gamma^5) u_{\nu_\alpha}^{(h)}(p)$$

$$\bar{u}_{\nu_\alpha}^{(h)}(p) \gamma^0 (1 - \gamma^5) u_{\nu_\alpha}^{(h)}(p)$$

$$= \text{Tr} \left[u_{\nu_\alpha}^{(h)}(p) \bar{u}_{\nu_\alpha}^{(h)}(p) \gamma^0 (1 - \gamma^5) \right]$$

$$= \text{Tr} \left[(\not{p} + m_{\nu_\alpha}) \left(\frac{1 + \gamma^5 \not{S}_h}{2} \right) \gamma^0 (1 - \gamma^5) \right]$$

(note that $\frac{1 + \gamma^5 \not{S}_h}{2}$ projects out helicity h)

$$= \text{Tr} \left[\frac{1}{2} \gamma^0 \not{p} - \frac{1}{2} m_{\nu_\alpha} \gamma^5 \not{S}_h \gamma^0 \gamma^5 \right]$$

$$= 4p^0 + \text{Tr} \left[-\frac{1}{2} m_{\nu_\alpha} \not{S}_h \gamma^0 \right]$$

$$= 4p^0 - 2h |\vec{p}|$$

$$= 4E \text{ if } h = -1$$

$$= m_{\nu_e}^2/E \text{ if } h = +1$$

$\left(S_h = h \left(\frac{|\vec{p}|}{m}, \frac{E}{m} \frac{\vec{p}}{|\vec{p}|} \right) \right)$

Hence

$$V_{\alpha}^{(-)} \approx V_{\alpha} \quad (\text{for } h = -1)$$

$$V_{\alpha}^{(+)} \approx V_{\alpha} \frac{m_{\nu_{\alpha}}^2}{4E^2} \quad (\text{for } h = +1)$$

One can also find out the potential energy of a $\bar{\nu}_{\alpha}$ (with helicity h) propagating through the medium:

$$\bar{V}_{\alpha}^{(h)} = \langle \bar{\nu}_{\alpha}(p, h) | \int d^3x \bar{\psi}(x) | \bar{\nu}_{\alpha}(p, h) \rangle$$

$$= \frac{-V_{\alpha}}{4E} \bar{V}_{\nu_{\alpha}}^{(h)}(p) \gamma^0 (1 - \gamma^5) V_{\nu_{\alpha}}^{(h)}(p)$$

$$= -V_{\alpha} \quad (\text{for } h = +1)$$

$$= -V_{\alpha} \frac{m_{\nu_{\alpha}}^2}{4E^2} \quad (\text{for } h = -1)$$

Therefore, the potential energy for a right-handed antineutrino has the same magnitude, but with an opposite sign with respect to the potential energy of a left-handed neutrino.

We are ready to consider neutrino oscillation in matter.

Consider a left-handed neutrino of flavor α produced at $t=0$:

$$|\nu_\alpha\rangle = \sum_k U_{\alpha k}^* |\nu_k\rangle$$

$$H = H_0 + H_I$$

H_0 is the Hamiltonian in vacuum, and H is the total Hamiltonian in matter

The mass eigenstates $|\nu_k\rangle$ correspond to the case in vacuum:

$$H_0 |\nu_k\rangle = E_k |\nu_k\rangle$$

while the flavor states $|\nu_\alpha\rangle$ are eigenstates of H_I :

$$H_I |\nu_\alpha\rangle = V_\alpha |\nu_\alpha\rangle$$

It is to be expected that neither $|\nu_k\rangle$ nor $|\nu_\alpha\rangle$ are the eigenstates of H .

The time evolution for $|\nu_\alpha(t)\rangle$ is

$$i \frac{d}{dt} |\nu_\alpha(t)\rangle = H |\nu_\alpha(t)\rangle$$

where $|\nu_\alpha(0)\rangle = |\nu_\alpha\rangle$

define the neutrino transition amplitude $\psi_{\alpha\beta}(t)$ as $\psi_{\alpha\beta}(t) = \langle \nu_\beta | \nu_\alpha(t) \rangle$

Then, the transition probability is

$$P_{\nu_\alpha \rightarrow \nu_\beta}(t) = |\psi_{\alpha\beta}(t)|^2$$

From $i \frac{d}{dt} |\nu_\alpha(t)\rangle = \hat{H} |\nu_\alpha(t)\rangle$

we have

$$\begin{aligned} \langle \nu_\beta | i \frac{d}{dt} |\nu_\alpha(t)\rangle &= \langle \nu_\beta | \hat{H} |\nu_\alpha(t)\rangle \\ &= \langle \nu_\beta | H_0 | \nu_\alpha(t)\rangle + \langle \nu_\beta | H_I | \nu_\alpha(t)\rangle \end{aligned}$$

The second term on the RHS is

$$\begin{aligned} \langle \nu_\beta | H_I | \nu_\alpha(t)\rangle &= \sum_{\eta} \langle \nu_\beta | H_I | \nu_\eta \rangle \langle \nu_\eta | \nu_\alpha(t)\rangle \\ &= \sum_{\eta} V_{\beta\eta} \delta_{\beta\eta} \psi_{\alpha\eta}(t) \quad (\eta \text{ is the flavor index}) \end{aligned}$$

The first term on the RHS is

$$\begin{aligned} \langle \nu_\beta | H_0 | \nu_\alpha(t)\rangle &= \sum_k \langle \nu_\beta | H_0 | \nu_k \rangle \langle \nu_k | \nu_\alpha(t)\rangle \\ &= \sum_{k, \eta} \langle \nu_\beta | H_0 | \nu_k \rangle \langle \nu_k | \nu_\eta \rangle \langle \nu_\eta | \nu_\alpha(t)\rangle \\ &= \sum_{k, \eta} U_{\beta k} E_k U_{\eta k}^* \psi_{\alpha\eta}(t) \quad (\eta \text{ is the mass eigenstate index, in vacuum}) \end{aligned}$$

Therefore, we obtain

$$i \frac{d}{dt} \psi_{\alpha\beta}(t) = \sum_{\eta} \left(\sum_k U_{\beta k} E_k U_{\eta k}^* + \delta_{\beta\eta} V_{\beta} \right) \psi_{\alpha\eta}(t)$$

Using $E_k \approx E + \frac{m_k^2}{2E}$, $t \approx x$, $p \approx E$,

$$V_\alpha = V_{cc} \delta_{\alpha e} + V_{\nu c}$$

We obtain

$$i \frac{d}{dx} \psi_{\alpha\beta}(x) = \left(p + \frac{m_1^2}{2E} + V_{NC} \right) \psi_{\alpha\beta}(x) + \sum_{\eta} \left(\sum_K U_{\beta K} \frac{\Delta m_{K1}^2}{2E} U_{\eta K}^* + \delta_{\beta e} \delta_{\eta e} V_{CC} \right) \psi_{\alpha\eta}(x)$$

The first term on the RHS is not relevant, since it corresponds to a phase in $\psi_{\alpha\beta}(x)$.

It can be eliminated by shifting the phase of $\psi_{\alpha\beta}(x)$, i.e., a shift of

$$\exp \left[-i \left(p + \frac{m_1^2}{2E} \right) x - i \int_0^x V_{NC}(x') dx' \right]$$

Therefore, the evolution of the transition amplitude is

$$i \frac{d}{dx} \psi_{\alpha\beta}(x) = \sum_{\eta} \left(\sum_K U_{\beta K} \frac{\Delta m_{K1}^2}{2E} U_{\eta K}^* + \delta_{\beta e} \delta_{\eta e} V_{CC} \right) \psi_{\alpha\eta}(x)$$

For a given α , the above equation contains $\beta = e, \mu, \tau$ parts. This is equivalent to a matrix equation

$$i \frac{d}{dx} \psi_{\alpha} = H_F \psi_{\alpha}$$

$$\text{Where } H_F = \frac{1}{2E} (U M^2 U^{\dagger} + A)$$

$$\psi_{\alpha} = \begin{pmatrix} \psi_{\alpha e} \\ \psi_{\alpha \mu} \\ \psi_{\alpha \tau} \end{pmatrix}; \quad M^2 = \begin{pmatrix} 0 & 0 & 0 \\ 0 & \Delta m_{21}^2 & 0 \\ 0 & 0 & \Delta m_{31}^2 \end{pmatrix}; \quad A = \begin{pmatrix} A_{CC} & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

$$A_{CC} = 2E V_{CC} = 2\sqrt{2} E G_F N_e$$

We consider the evolution for $i \frac{d}{dx} \psi_\alpha = H_F \psi_\alpha$
for the two-neutrino mixing case:

$$\nu_e = \nu_\mu \quad (\nu_1 \text{ and } \nu_2)$$

$$\psi_\alpha = \begin{pmatrix} \psi_{\alpha e} \\ \psi_{\alpha \mu} \end{pmatrix}; \quad M^2 = \begin{pmatrix} 0 & 0 \\ 0 & \Delta m^2 \end{pmatrix}; \quad A = \begin{pmatrix} A_{cc} & 0 \\ 0 & 0 \end{pmatrix}$$

$$A_{cc} = 2E V_{cc} = 2\sqrt{2} E G_F N_e$$

$$H_F = \frac{1}{2E} (U M^2 U^\dagger + A)$$

One can make M^2 and A more symmetric
by adding diagonal matrices

$$M^2 = \begin{pmatrix} 0 & 0 \\ 0 & \Delta m^2 \end{pmatrix} \longrightarrow \begin{pmatrix} 0 & 0 \\ 0 & \Delta m^2 \end{pmatrix} + \begin{pmatrix} -\frac{\Delta m^2}{2} & 0 \\ 0 & -\frac{\Delta m^2}{2} \end{pmatrix}$$

$$= \begin{pmatrix} -\frac{\Delta m^2}{2} & 0 \\ 0 & \frac{\Delta m^2}{2} \end{pmatrix} = (M^2)'$$

$$A = \begin{pmatrix} A_{cc} & 0 \\ 0 & 0 \end{pmatrix} \longrightarrow \begin{pmatrix} A_{cc} & 0 \\ 0 & 0 \end{pmatrix} + \begin{pmatrix} -\frac{A_{cc}}{2} & 0 \\ 0 & -\frac{A_{cc}}{2} \end{pmatrix}$$

$$= \begin{pmatrix} \frac{A_{cc}}{2} & 0 \\ 0 & -\frac{A_{cc}}{2} \end{pmatrix} = A'$$

$$H_F \longrightarrow H_F + \left(-\frac{\Delta m^2}{2} - \frac{A_{cc}}{2} \right) I = H_F'$$

Therefore H_F' and H_F differ by a term
proportional to a unit matrix aI ,

where $a = \left(+\frac{\Delta m^2}{2} + \frac{A_{cc}}{2} \right) / 2E$. a can depend on x
since A_{cc} can depend on x (due to $N_e(x)$)

$$\text{Let } \psi'_\alpha = \exp\left\{i \int_0^x a(x') dx'\right\} \psi_\alpha$$

then

$$i \frac{d}{dx} \psi'_\alpha = \exp\left\{i \int_0^x a(x') dx'\right\} i \frac{d}{dx} \psi_\alpha - a(x) \exp\left\{i \int_0^x a(x') dx'\right\} \psi_\alpha$$

$$= \exp\left\{i \int_0^x a(x') dx'\right\} H_F \psi_\alpha$$

$$- a(x) \exp\left\{i \int_0^x a(x') dx'\right\} \psi_\alpha$$

$$= [H_F - a(x) I] \psi'_\alpha = H'_F \psi'_\alpha$$

Hence ψ'_α satisfies the same differential equation as ψ_α , provided that $H_F \rightarrow H'_F = H_F - a(x) I$

Since ψ'_α and ψ_α only differ by a phase factor, the transition probability, $|\psi_\alpha|^2$, remains the same for $|\psi'_\alpha|^2$.

Therefore, we are free to change M^2 to $\begin{pmatrix} -\frac{\Delta m^2}{2} & 0 \\ 0 & \frac{\Delta m^2}{2} \end{pmatrix}$

and A to $\begin{pmatrix} \frac{Acc}{2} & 0 \\ 0 & -\frac{Acc}{2} \end{pmatrix}$. Note the $U = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix}$.

$$UMU^\dagger + A = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} -\frac{\Delta m^2}{2} & 0 \\ 0 & \frac{\Delta m^2}{2} \end{pmatrix} \begin{pmatrix} \cos\theta & -\sin\theta \\ \sin\theta & \cos\theta \end{pmatrix} + \begin{pmatrix} \frac{Acc}{2} & 0 \\ 0 & -\frac{Acc}{2} \end{pmatrix}$$

$$= \frac{1}{2} \begin{pmatrix} -\Delta m^2 \cos 2\theta + Acc & \Delta m^2 \sin 2\theta \\ \Delta m^2 \sin 2\theta & \Delta m^2 \cos 2\theta - Acc \end{pmatrix}$$

The differential equation becomes

$$i \frac{d}{dx} \begin{pmatrix} \psi_{ee} \\ \psi_{\mu e} \end{pmatrix} = \frac{1}{4E} \begin{pmatrix} -\Delta m^2 \cos 2\theta + A_{cc} & \Delta m^2 \sin 2\theta \\ \Delta m^2 \sin 2\theta & \Delta m^2 \cos 2\theta - A_{cc} \end{pmatrix} \begin{pmatrix} \psi_{ee} \\ \psi_{\mu e} \end{pmatrix}$$

note the θ is the mixing angle in vacuum.

$$\nu_e = \cos \theta \nu_1 + \sin \theta \nu_2 ; \quad \nu_\mu = -\sin \theta \nu_1 + \cos \theta \nu_2$$

$$\Delta m^2 = \Delta m_{21}^2 = m_2^2 - m_1^2, \text{ which is mass-squared difference in vacuum.}$$

$$\text{For } H_F = \frac{1}{4E} \begin{pmatrix} -\Delta m^2 \cos 2\theta + A_{cc} & \Delta m^2 \sin 2\theta \\ \Delta m^2 \sin 2\theta & \Delta m^2 \cos 2\theta - A_{cc} \end{pmatrix}$$

we would like to diagonalize it

$$U_M^\top H_F U_M = H_M, \text{ where } H_M \text{ is diagonal}$$

$$\text{write } U_M = \begin{pmatrix} \cos \theta_M & \sin \theta_M \\ -\sin \theta_M & \cos \theta_M \end{pmatrix}$$

$$\text{and } H_M = \frac{1}{4E} \begin{pmatrix} -\Delta m_M^2 & 0 \\ 0 & \Delta m_M^2 \end{pmatrix}$$

It is straightforward to show that

$$\Delta m_M^2 = \left[(\Delta m^2 \cos 2\theta - A_{cc})^2 + (\Delta m^2 \sin 2\theta)^2 \right]^{1/2}$$

$$A_{cc} = 2\sqrt{2} E G_F N_e(x)$$

The minimal value of Δm_M^2 is $\Delta m^2 \sin 2\theta$, which occurs at $\Delta m^2 \cos 2\theta = A_{cc}$

It is simple to show

$$\tan 2\theta_M = \frac{\Delta m^2 \sin 2\theta}{\Delta m^2 \cos 2\theta - A_{cc}} = \frac{\tan 2\theta}{1 - \frac{A_{cc}}{\Delta m^2 \cos 2\theta}}$$

When $\Delta m^2 \cos 2\theta - A_{cc} = 0$, $\theta_M = \pi/4$, maximal mixing

Therefore, even with a tiny value of θ , one can obtain maximal mixing $\theta_M = \pi/4$, if the electron density is such that $A_{cc}^R = \Delta m^2 \cos 2\theta$ (we call this 'Resonance', hence the notation R)

Since $A_{cc} = 2\sqrt{2} E G_F N_e$, the 'resonance' electron density occurs at

$$N_e^R = \frac{\Delta m^2 \cos 2\theta}{2\sqrt{2} E G_F}$$

* This density can be reached if Δm^2 is small enough, or if E is large enough. For example, for $\Delta m^2 = 10^{-5} \text{ eV}^2$, $\theta = 0.1^\circ$, $E = 1 \text{ MeV}$,

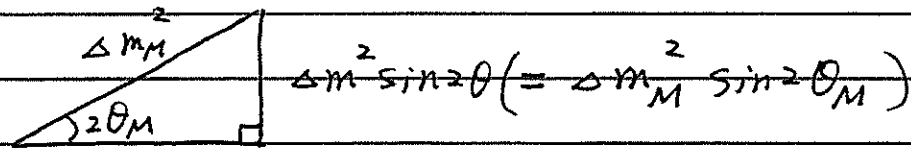
we obtain $N_e^R = 65.6 N_A / \text{cm}^3$

where N_A is the Avogadro number.

This density is reachable in SUN.

* Note that if $\theta = \pi/4$, then the resonant electron density is zero, and maximal mixing occurs in vacuum.

One can also show graphically :



$$\Delta m^2 \cos 2\theta - Acc$$

$$(\equiv \Delta m_M^2 \cos 2\theta_M)$$

(From $\tan 2\theta_M = \frac{\Delta m^2 \sin 2\theta}{\Delta m^2 \cos 2\theta - Acc}$, one obtains

$$\sin 2\theta_M = \frac{\Delta m^2 \sin 2\theta}{[(\Delta m^2 \cos 2\theta - Acc)^2 + (\Delta m^2 \sin 2\theta)^2]^{1/2}} = \frac{\Delta m^2 \sin 2\theta}{\Delta m_M^2}$$

and

$$\cos 2\theta_M = \frac{\Delta m^2 \cos 2\theta - Acc}{[(\Delta m^2 \cos 2\theta - Acc)^2 + (\Delta m^2 \sin 2\theta)^2]^{1/2}} = \frac{\Delta m^2 \cos 2\theta - Acc}{\Delta m_M^2}$$

Note that $\Delta m^2 \sin 2\theta$ part of this triangle is x -independent (or density independent), while the $\Delta m^2 \cos 2\theta - Acc$ part of the triangle depends on the local electron density N_e .

One can see that the HF matrix can be written as

$$HF = \frac{1}{4E} \begin{pmatrix} -\Delta m^2 \cos 2\theta + Acc & \Delta m^2 \sin 2\theta \\ \Delta m^2 \sin 2\theta & \Delta m^2 \cos 2\theta - Acc \end{pmatrix} = \frac{1}{4E} \begin{pmatrix} \Delta m_M^2 \cos 2\theta_M & \Delta m_M^2 \sin 2\theta_M \\ \Delta m_M^2 \sin 2\theta_M & \Delta m_M^2 \cos 2\theta_M \end{pmatrix}$$

After diagonalizing H_F , how would the evolution equation, $i \frac{d}{dx} \psi_\alpha = H_F \psi_\alpha$, be modified?

We have

$$i \frac{d}{dx} \begin{pmatrix} \psi_{ee} \\ \psi_{em} \end{pmatrix} = U_M H_M U_M^\dagger \begin{pmatrix} \psi_{ee} \\ \psi_{em} \end{pmatrix}$$

We can define

$$\Phi_e = \begin{pmatrix} \phi_{e1} \\ \phi_{e2} \end{pmatrix} = U_M^\dagger \begin{pmatrix} \psi_{ee} \\ \psi_{em} \end{pmatrix} = U_M^\dagger \psi_e ; \psi_e = \begin{pmatrix} \psi_{ee} \\ \psi_{em} \end{pmatrix}$$

ϕ_{e1}, ϕ_{e2} correspond to the transition amplitude for $|e\rangle$ at $t=0$ to transform into $|v_1\rangle_M, |v_2\rangle_M$, which are the mass eigenstates in matter (for a given N_e electron density) later at time t .

The evolution equation can be expressed in terms of Φ_e as (since $\psi_e = U_M \Phi_e$)

$$i \frac{d}{dx} \psi_e = U_M H_M \Phi_e$$

or

$$U_M^\dagger i \frac{d}{dx} (U_M \Phi_e) = H_M \Phi_e \quad \left(H_M = \frac{1}{4E} \begin{pmatrix} -\Delta m_{21}^2 & 0 \\ 0 & \Delta m_{21}^2 \end{pmatrix} \right)$$

The LHS becomes

$$\begin{aligned} & U_M^\dagger i \frac{d}{dx} (U_M) \begin{pmatrix} \phi_{e1} \\ \phi_{e2} \end{pmatrix} + i U_M^\dagger U_M \frac{d}{dx} \begin{pmatrix} \phi_{e1} \\ \phi_{e2} \end{pmatrix} \\ &= i \begin{pmatrix} \cos \theta_M & -\sin \theta_M \\ \sin \theta_M & \cos \theta_M \end{pmatrix} \frac{d}{dx} \begin{pmatrix} \cos \theta_M & \sin \theta_M \\ -\sin \theta_M & \cos \theta_M \end{pmatrix} \begin{pmatrix} \phi_{e1} \\ \phi_{e2} \end{pmatrix} + i \frac{d}{dx} \begin{pmatrix} \phi_{e1} \\ \phi_{e2} \end{pmatrix} \end{aligned}$$

$$i U_M^\dagger \frac{d}{dx} [U_M]$$

$$= i \begin{pmatrix} \cos \theta_M & -\sin \theta_M \\ \sin \theta_M & \cos \theta_M \end{pmatrix} \begin{pmatrix} -\sin \theta_M \theta_M' & \cos \theta_M \theta_M' \\ -\cos \theta_M \theta_M' & -\sin \theta_M \theta_M' \end{pmatrix}$$

$$= i \theta_M' \begin{pmatrix} 0 & +1 \\ -1 & 0 \end{pmatrix} \quad \text{Where } \theta_M' = \frac{d\theta_M}{dx}$$

Hence the evolution equation becomes

$$i \frac{d}{dx} \begin{pmatrix} \phi_{e1} \\ \phi_{e2} \end{pmatrix} = \frac{1}{4E} \begin{pmatrix} -\Delta m_M^2 & -4Ei \theta_M' \\ 4Ei \theta_M' & \Delta m_M^2 \end{pmatrix} \begin{pmatrix} \phi_{e1} \\ \phi_{e2} \end{pmatrix}$$

$$\frac{d\theta_M}{dx} = \frac{1}{2} \frac{\sin 2\theta_M}{\Delta m_M^2} \frac{dA_{cc}}{dx}$$

1) The simplest case is a constant N_e , i.e. $\theta_M' = 0$, corresponding to propagation in vacuum or uniform medium:

ϕ_{e1} and ϕ_{e2} are decoupled, Δm_M^2 is constant, the evolution for ϕ_{e1} , ϕ_{e2} just lead to

a change in phase $\phi_{e_i}(x) = \phi_{e_i}(0) \exp\left(\frac{i \Delta m_M^2}{4E} x\right)$

$$\begin{pmatrix} \phi_{e1}(0) \\ \phi_{e2}(0) \end{pmatrix} = \begin{pmatrix} \cos \theta_M^{(i)} & -\sin \theta_M^{(i)} \\ \sin \theta_M^{(i)} & \cos \theta_M^{(i)} \end{pmatrix} \begin{pmatrix} 1 \\ 0 \end{pmatrix} = \begin{pmatrix} \cos \theta_M^{(i)} \\ \sin \theta_M^{(i)} \end{pmatrix}$$

$$P_{\nu_e \rightarrow \nu_\mu}(x) = \sin^2 2\theta_M \sin^2 \left(\frac{\Delta m_M^2 x}{4E} \right)$$

2) Another case, so-called adiabatic evolution, is for a non-zero θ_M' , but θ_M' is negligible compared to the diagonal terms:

$$\gamma = \frac{\Delta m_M^2}{4E|\theta_M'|} \gg 1$$

where γ is the "adiabaticity" parameter

Note that ϕ_{e1} and ϕ_{e2} are decoupled in this case, however, Δm_M^2 depends on x , hence the solutions for $\phi_{e1}(x)$ and $\phi_{e2}(x)$ are

$$\phi_{e1}(x) = \exp\left[i \int_0^x \frac{\Delta m_M^2(x')}{4E} dx'\right] \phi_{e1}(0)$$

$$\phi_{e2}(x) = \exp\left[-i \int_0^x \frac{\Delta m_M^2(x')}{4E} dx'\right] \phi_{e2}(0)$$

$$\text{Let } \int_0^x \frac{\Delta m_M^2(x')}{4E} dx' = A$$

then, the survival amplitude $\phi_{ee}(x)$ is

$$\phi_{ee}(x) = \cos \theta_M^{(+)} \phi_{e1}(x) + \sin \theta_M^{(+)} \phi_{e2}(x)$$

$$= \cos \theta_M^{(+)} \cos \theta_M^{(i)} \exp(iA) + \sin \theta_M^{(+)} \sin \theta_M^{(i)} \exp(-iA)$$

and the survival probability is

$$|\phi_{ee}(x)|^2 = P_{\nu_e \rightarrow \nu_e}(x) = \frac{1}{2} + \frac{1}{2} \cos 2\theta_M^{(i)} \cos 2\theta_M^{(+)} + \frac{1}{2} \sin 2\theta_M^{(i)} \sin 2\theta_M^{(+)} \times \cos(2A)$$

(If $\theta_M^{(i)} = \theta_M^{(f)} = \theta$ (vacuum)

then

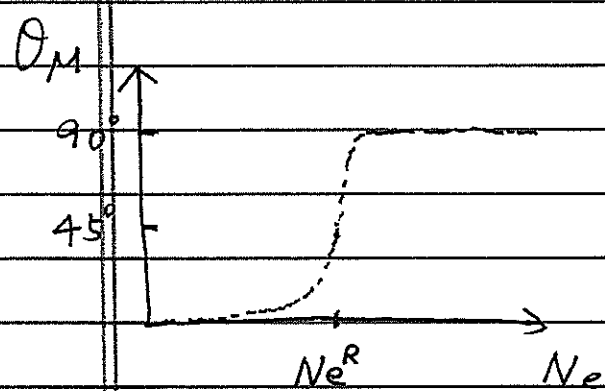
$$\begin{aligned}
 P_{\nu_e \rightarrow \nu_e} &= \frac{1}{2} + \frac{1}{2} \cos^2 2\theta + \frac{1}{2} \sin^2 2\theta \cos(2A) \\
 &= \frac{1}{2} + \frac{1}{2} (1 - \sin^2 2\theta) + \frac{1}{2} \sin^2 2\theta \cos(2A) \\
 &= 1 - \frac{1}{2} \sin^2 2\theta (1 - \cos 2A) \\
 &= 1 - \frac{1}{2} \sin^2 2\theta (2 \sin^2 A) = 1 - \sin^2 2\theta \sin^2 \left(\frac{\Delta m^2 L}{4E} \right)
 \end{aligned}$$

as expected.

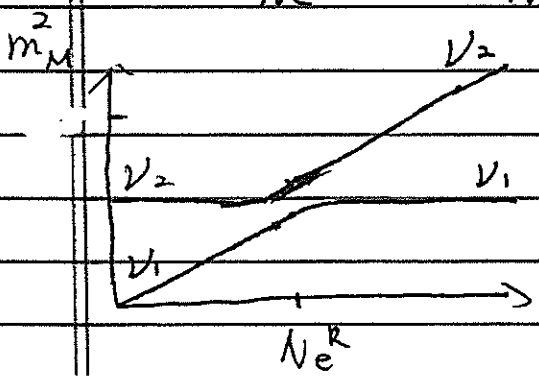
In general, the 'adiabatic' transition implied an averaged survival probability

($\overline{P}_{\nu_e \rightarrow \nu_e}^{\text{adiabatic}} = \frac{1}{2} + \frac{1}{2} \cos 2\theta_M^{(i)} \cos 2\theta_M^{(f)}$

Consider a ν_e produced near the center of sun, where $N_e > N_e^R$ (and small θ)



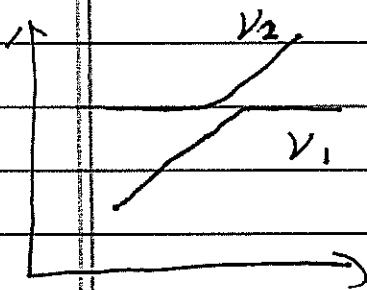
$\theta_M^{(I)} \cong 90^\circ$
 $\tan^2 \theta_M = \frac{\Delta m^2 \sin^2 \theta}{\Delta m^2 \cos^2 \theta - A_{cc}}$
 $\nu_e = \cos \theta_M \nu_1 + \sin \theta_M \nu_2$
 $\theta_M^{(I)} \cong 90^\circ \Rightarrow \nu_e \sim \nu_2 \text{ at } t=0$



Adiabaticity implies that no transition (coupling) between ν_1 and ν_2
 $\theta_M^{(f)} = \theta \rightarrow 0 \Rightarrow \nu_2 \sim \nu_\mu$
 $P_{\nu_e \rightarrow \nu_e} \cong \frac{1}{2} + \frac{1}{2} \cos(180^\circ) \cos(0) = 0$
 $P_{\nu_e \rightarrow \nu_\mu} \cong 1$

3) In the non-adiabatic case, when γ near the resonance is small, then one can either find out the solution numerically (solve the coupled differential equation), or use an approximation as follows:

since γ at the region away from the resonance is large (adiabatic), we only need to consider the situation when ν_e crosses the resonance.



The probability for $\nu_2 \rightarrow \nu_1$ at the resonance-crossing is

$$P_c = P_{\nu_2 \rightarrow \nu_1} = P_{\nu_1 \rightarrow \nu_2}$$

The probability for $\nu_2 \rightarrow \nu_2$ while

crossing the resonance is

$$1 - P_c = P_{\nu_2 \rightarrow \nu_2} = P_{\nu_1 \rightarrow \nu_1}$$

$$P_c = \frac{\exp\left[-\frac{\pi}{2} \gamma_R F\right] - \exp\left[-\frac{\pi}{2} \gamma_R \frac{F}{\sin^2 \theta}\right]}{1 - \exp\left[-\frac{\pi}{2} \gamma_R \frac{F}{\sin^2 \theta}\right]}$$

F depends on the shape of $N_e(x)$, and $F \approx 1$

For $\gamma_R \rightarrow 0$, we have $P_c \rightarrow \cos^2 \theta$

and $\overline{P_{\nu_e \rightarrow \nu_e}} (\gamma_R \rightarrow 0) \approx 1 - \frac{1}{2} \sin^2 \theta$

just like in vacuum. For abrupt density change at the resonance, ν_e maintains its flavor, and oscillate like ν_e afterwards.

$$\overline{P_{\nu_e \rightarrow \nu_e}} = \frac{1}{2} + \left(\frac{1}{2} - P_c\right) \cos^2 \theta_M^{(i)} \cos^2 \theta$$

1) In vacuum: $\theta_M^{(i)} = \theta$, $P_c = 0$

$$\overline{P_{\nu_e \rightarrow \nu_e}} = \frac{1}{2} + \frac{1}{2} \cos^2 2\theta = 1 - \frac{1}{2} \sin^2 \theta$$

2) Adiabatic: $P_c = 0$

$$\overline{P_{\nu_e \rightarrow \nu_e}} = \frac{1}{2} + \frac{1}{2} \cos^2 \theta_M^{(i)} \cos^2 \theta$$

a) $N_e^{(i)} \gg N_e^R$, $\theta_M^{(i)} = \pi/2$

$$\overline{P_{\nu_e \rightarrow \nu_e}} = \frac{1}{2} - \frac{1}{2} \cos^2 \theta$$

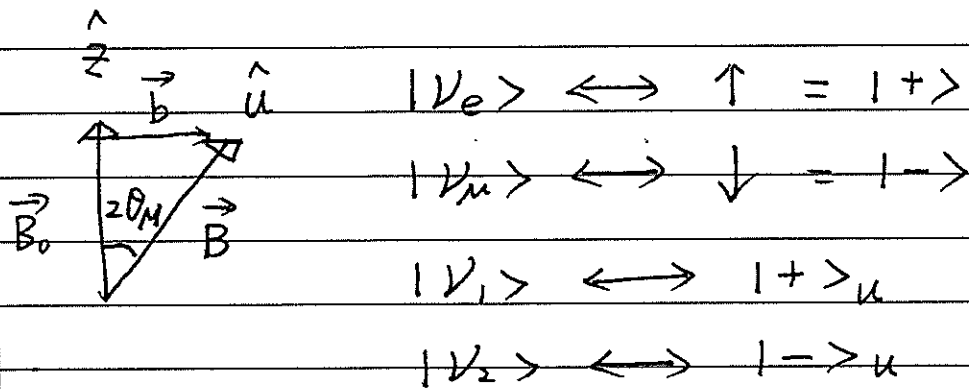
b) $N_e^{(i)} \ll N_e^R$, $\theta_M^{(i)} = \theta$

$$\begin{aligned} \overline{P_{\nu_e \rightarrow \nu_e}} &= \frac{1}{2} + \frac{1}{2} \cos^2 2\theta = \frac{1}{2} + \frac{1}{2} (1 - \sin^2 2\theta) \\ &= 1 - \frac{1}{2} \sin^2 2\theta \end{aligned}$$

3) Abrupt: $P_c = \cos^2 \theta$

$$\overline{P_{\nu_e \rightarrow \nu_e}} = 1 - \frac{1}{2} \sin^2 2\theta$$

In QM, there is a well known analogy between mixing of a two-level system and the behavior of spin- $\frac{1}{2}$ particle in magnetic field.



At $t=0$, $|V_e\rangle$ is like a spin- $\frac{1}{2}$ particle pointing along the \hat{z} axis with the probability of $|1+\rangle$ being $\cos^2(\frac{\theta}{2}) = 1$ and $|1-\rangle$ being $\sin^2(\frac{\theta}{2}) = 0$

$$H = -\gamma \vec{B} \cdot \vec{S} = -\frac{\gamma \hbar}{2} \begin{pmatrix} B_z & B_x - iB_y \\ B_x + iB_y & -B_z \end{pmatrix}$$

Let $B_z = |\vec{B}_0|$, $B_y = 0$, $B_x = |\vec{b}|$

then

$$H \sim \begin{pmatrix} B_0 & B_x \\ B_x & -B_0 \end{pmatrix}$$

which is the same form as

$$\frac{1}{4E} \begin{pmatrix} -\Delta m^2 \cos 2\theta + A_{cc} & \Delta m^2 \sin 2\theta \\ \Delta m^2 \sin 2\theta & \Delta m^2 \cos 2\theta - A_{cc} \end{pmatrix}$$

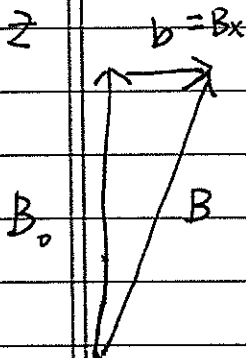
$$| \psi_e \rangle \leftrightarrow \uparrow = | + \rangle$$

$$| \psi_m \rangle \leftrightarrow \downarrow = | - \rangle$$

$H = -\gamma \vec{B} \cdot \vec{S}$, hence \uparrow and \downarrow are eigenstates of H when \vec{B} is along the \hat{z} axis

$$H_0 = -\gamma \vec{B}_0 \cdot \vec{S} = \begin{pmatrix} -\frac{\gamma \hbar}{2} B_0 & 0 \\ 0 & +\frac{\gamma \hbar}{2} B_0 \end{pmatrix}$$

\hat{z}



eigenstate for

$$H = H_0 + H_I = -\gamma (\vec{B}_0 + \vec{b}) \cdot \vec{S}$$

$$H = +\frac{\gamma \hbar}{2} \begin{pmatrix} -B_0 & b_x - i b_y \\ b_x + i b_y & +B_0 \end{pmatrix}$$

$$\vec{b} = B_x \hat{x}$$

$$H_I = \frac{\gamma \hbar}{2} \begin{pmatrix} -B_0 & B_x \\ B_x & +B_0 \end{pmatrix}$$

eigenval

$$H_{diag} = U^T H U$$

$$U = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix}$$

$$H_{diag} = +\frac{\gamma \hbar}{2} \begin{pmatrix} -B & 0 \\ 0 & B \end{pmatrix} \quad B = (B_0^2 + B_x^2)^{1/2}$$

$$\tan 2\theta = \frac{B_x}{B_0}$$

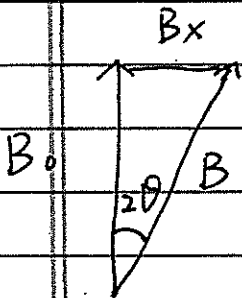
$$B_x = B_0 \tan^2 \theta, \quad B = B_0 (1 + \tan^2 \theta)^{1/2} \\ = B_0 / \cos 2\theta$$

$$\text{or } B_0 = B \cos 2\theta$$

$$B_x = B \sin 2\theta$$

Hence

$$H = \frac{\gamma \hbar}{2} \begin{pmatrix} -B \cos 2\theta & B \sin 2\theta \\ B \sin 2\theta & B \cos 2\theta \end{pmatrix}$$



For a spin- $1/2$ particle

The eigenstate becomes

$|+\rangle_u$ and $|-\rangle_u$ where

\hat{u} points along the direction of \hat{B}

$$|+\rangle_u = \cos \theta |+\rangle + \sin \theta |-\rangle$$

$$|-\rangle_u = -\sin \theta |+\rangle + \cos \theta |-\rangle$$

$$\text{if } \theta = 0, \quad |+\rangle_u = |+\rangle$$

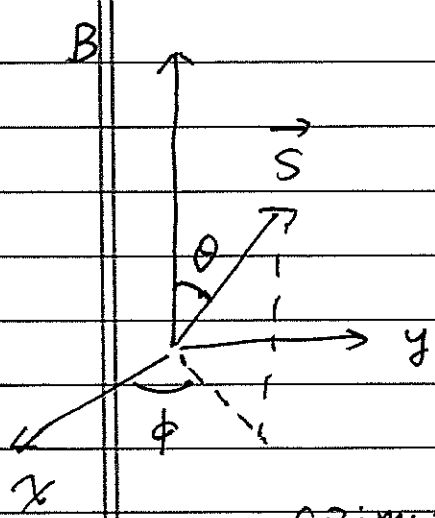
$$\text{if } \theta = \pi/2, \quad |+\rangle_u = |-\rangle$$

$$\text{if } \theta = \pi/4, \quad |+\rangle_u = \frac{1}{\sqrt{2}} (|+\rangle + |-\rangle)$$

$$|\langle + | + \rangle_u|^2 = \cos^2 \theta; \quad |\langle - | + \rangle_u|^2 = \sin^2 \theta$$

Precession Larmor frequency $\omega = \gamma B$

(Precession of spin-1/2 in B field



$$H = -\gamma \vec{B} \cdot \vec{S}$$

$$\vec{S} = \begin{pmatrix} \cos \frac{\theta}{2} e^{-i\phi/2} \\ \sin \frac{\theta}{2} e^{i\phi/2} \end{pmatrix}; \vec{S} = \frac{\hbar}{2} \hat{S}$$

θ, ϕ are the polar and azimuthal angles

(
$$\psi(t=0) = \begin{pmatrix} \cos \frac{\theta}{2} e^{-i\phi/2} \\ \sin \frac{\theta}{2} e^{i\phi/2} \end{pmatrix}$$

$$\psi(t) = e^{-\frac{iHt}{\hbar}} \begin{pmatrix} \cos \frac{\theta}{2} e^{-i\phi/2} \\ \sin \frac{\theta}{2} e^{i\phi/2} \end{pmatrix}$$

$$e^{-\frac{iHt}{\hbar}} = \exp \left\{ -\frac{i}{\hbar} t \begin{pmatrix} -\gamma B \frac{\hbar}{2} & 0 \\ 0 & \gamma B \frac{\hbar}{2} \end{pmatrix} \right\}$$

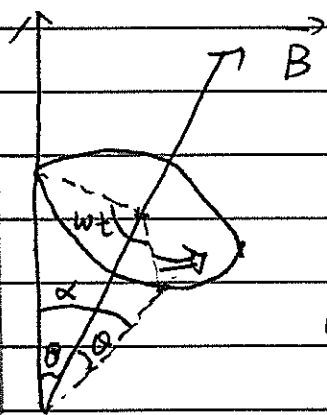
$$= \exp \begin{bmatrix} i \frac{\gamma B}{2} t & 0 \\ 0 & -i \frac{\gamma B}{2} t \end{bmatrix} = \begin{pmatrix} \exp i(\frac{\gamma B}{2} t) & 0 \\ 0 & \exp -i(\frac{\gamma B}{2} t) \end{pmatrix}$$

(
$$\psi(t) = \begin{pmatrix} \cos \frac{\theta}{2} e^{-i(\phi - \gamma B t)} \\ \sin \frac{\theta}{2} e^{-i(\phi - \gamma B t)} \end{pmatrix}$$
 precession with $\omega = \gamma B$

(After adding \vec{b} to \vec{B} , the diagonal H' becomes

$$\frac{1}{4E} \begin{pmatrix} -\Delta m_M^2 & 0 \\ 0 & \Delta m_M^2 \end{pmatrix}$$

and the rotating angle is θ_M , and the new axis points along \hat{u} , making an angle of $2\theta_M$ with respect to \hat{z} .



precession frequency ω is $\frac{\Delta m_M^2}{4E}$

$$\cos \alpha = \cos^2 \theta + \sin^2 \theta \cos \omega t$$

and the probability to find a $|e\rangle$ at time t is ($\theta = 2\theta_M$)

$$P_{\nu_e \rightarrow \nu_e} = \cos^2 \frac{\alpha}{2} = \frac{1 + \cos \alpha}{2}$$

$$= \frac{1 + \cos^2 \theta + \sin^2 \theta \cos \omega t}{2} = \frac{1 + 1 - \sin^2(2\theta_M) + \sin^2 2\theta_M \cos \omega t}{2}$$

$$= 1 - \sin^2(2\theta_M) \cdot \sin^2\left(\frac{\Delta m_M^2 t}{4E}\right)$$

$$| \nu_e \rangle = | \nu_+ \rangle, \quad | \nu_\mu \rangle = | \nu_- \rangle$$

$$H = \frac{1}{4E} \begin{pmatrix} -\Delta m_M^2 \cos 2\theta_M & \Delta m_M^2 \sin 2\theta_M \\ \Delta m_M^2 \sin 2\theta_M & \Delta m_M^2 \cos 2\theta_M \end{pmatrix}$$

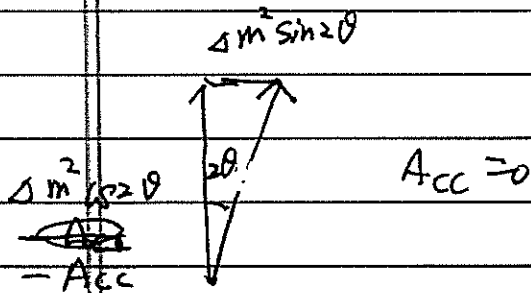
$$= \frac{1}{4E} \begin{pmatrix} -\Delta m^2 \cos 2\theta + A_{cc} & \Delta m^2 \sin 2\theta \\ \Delta m^2 \sin 2\theta & \Delta m^2 \cos 2\theta - A_{cc} \end{pmatrix}$$

Vacuum mass eigenstate

If $\theta_M = 0 = \theta$

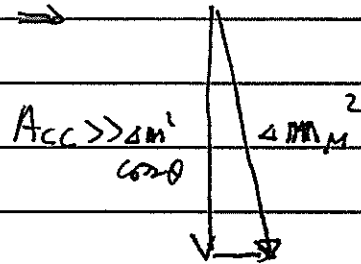
become

mass eigenstate = flavor eigenstate



In matter

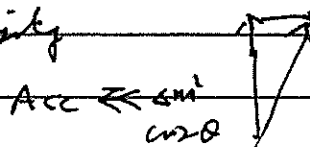
1) high density \Rightarrow



2) resonance \Rightarrow

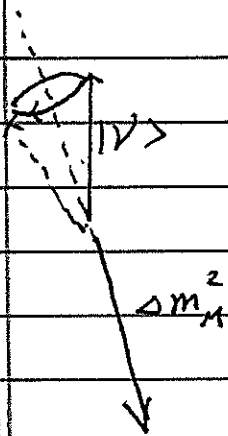
$$A_{cc} = \Delta m^2 \cos 2\theta \rightarrow \Delta m_M^2$$

3) low density



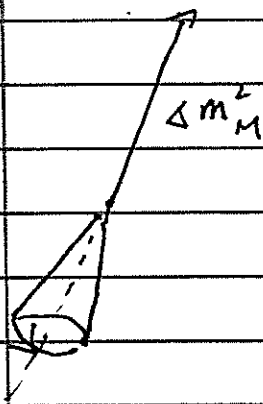
Now, consider ν_e produced at a density which is well above the resonance density, i.e.

$$N_e \gg N_e^R, \quad \Delta m_{\nu}^2$$



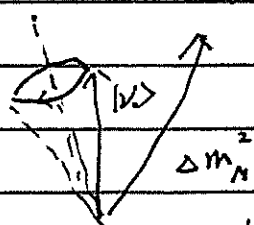
$|\nu_e\rangle_t$ is roughly $|\nu_2\rangle$, and precess in a small cone along the axis of the mass eigenstate

For adiabatic crossing of the resonance, $|\nu\rangle_t$ precess along an axis whose direction continues to change. Eventually, after crossing the resonance, it becomes



therefore, $|\nu\rangle$ becomes largely $|\nu_1\rangle$ for adiabatic transition

If the density changes too rapidly ($\dot{N}_e \gg \frac{4E}{\Delta m_{\nu}^2}$), then we have



Since $|\nu\rangle_t$ can not follow the rapid change of mass eigenstates.

Hence $|\nu\rangle$ remains as $|\nu_e\rangle$ after crossing the resonance.