

## PHYS 515: Homework Set 3

**Due Date:** Tuesday February 25, 2020, at the beginning of class.

**Topic:** Graduate General Relativity 1

1. Carroll 1.8

If  $\partial_\nu T^{\mu\nu} = Q^\mu$ , what physically does the spatial vector  $Q^i$  represent? Use the dust energy momentum tensor to make your case.

2. Carroll 1.9

For a system of discrete point particles the energy-momentum tensor takes the form

$$T_{\mu\nu} = \sum_a \frac{p_\mu^{(a)} p_\nu^{(a)}}{p^{0(a)}} \delta^{(3)}(\mathbf{x} - \mathbf{x}^{(a)}),$$

where the index  $a$  labels the different particles. Show that, for a dense collection of particles with isotropically distributed velocities, we can smooth over the individual particle worldlines to obtain the perfect-fluid energy-momentum tensor (1.114).

3. Schutz 4.24

Astronomical observations of the brightness of objects are measurements of the flux of radiation  $T^{01}$  from the object at Earth. This problem calculates how that flux depends on the relative velocity of the object and Earth.

a) Show that, in the rest frame  $\mathcal{O}$  of a star of constant luminosity  $L$  (total energy radiated per second), the stress-energy tensor of the radiation from the star at the event  $(t, x, 0, 0)$  has components  $T^{00} = T^{0x} = T^{x0} = T^{xx} = L/(4\pi x^2)$ . The star sits at the origin.

b) Let  $\vec{X}$  be the null vector that separates the events of emission and reception of the radiation. Show that  $\vec{X} \rightarrow_{\mathcal{O}} (x, x, 0, 0)$  for radiation observed at the event  $(x, x, 0, 0)$ . Show that the stress-energy tensor of (a) has the frame-invariant form

$$\mathbf{T} = \frac{L}{4\pi} \frac{\vec{X} \otimes \vec{X}}{(\vec{U}_s \cdot \vec{X})^4},$$

where  $\vec{U}_s$  is the star's four-velocity,  $\vec{U}_s \rightarrow_{\mathcal{O}} (1, 0, 0, 0)$ .

c) Let the Earth-bound observer  $\bar{\mathcal{O}}$ , traveling with speed  $v$  away from the star in the  $x$  direction, measure the same radiation, again with the star on the  $\bar{x}$  axis. Let  $\vec{X} \rightarrow_{\bar{\mathcal{O}}} (R, R, 0, 0)$  and find  $R$  as a function of  $x$ . Express  $T^{\bar{0}\bar{x}}$  in terms of  $R$ . Explain why  $R$  and  $T^{\bar{0}\bar{x}}$  depend as they do on  $v$ .

4. Schutz 4.25

Maxwell's equations for the electric and magnetic fields in vacuum,  $\mathbf{E}$  and  $\mathbf{B}$ , in three-vector notation are

$$\begin{aligned} \nabla \times \mathbf{B} - \frac{\partial}{\partial t} \mathbf{E} &= 4\pi \mathbf{J}, \\ \nabla \times \mathbf{E} + \frac{\partial}{\partial t} \mathbf{B} &= 0, \\ \nabla \cdot \mathbf{E} &= 4\pi \rho, \\ \nabla \cdot \mathbf{B} &= 0, \end{aligned}$$

in units where  $\mu_0 = \epsilon_0 = c = 1$ . (Here  $\rho$  is the density of the electric charge and  $\mathbf{J}$  the current density.)

a) An *antisymmetric*  $\binom{2}{0}$  tensor  $\mathbf{F}$  can be defined on spacetime by the equations  $F^{0i} = E^i$ ,  $F^{xy} = B^z$ ,  $F^{yz} = B^x$ ,  $F^{zx} = B^y$ . Find from this definition all other components  $F^{\mu\nu}$  in this frame and write them down in matrix in a matrix.

b) A rotation by an angle  $\theta$  about the  $z$  axis is one kind of Lorentz transformation, with the matrix

$$\Lambda^{\beta'}_{\alpha} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & \cos \theta & -\sin \theta & 0 \\ 0 & \sin \theta & \cos \theta & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} .$$

Show that the new components of  $\mathbf{F}$ ,

$$F^{\alpha'\beta'} = \Lambda^{\alpha'}_{\mu} \Lambda^{\beta'}_{\nu} F^{\mu\nu} ,$$

define new electric and magnetic three-vector components (by the rule given in (a)) that are just the same as the components of the old  $\mathbf{E}$  and  $\mathbf{B}$  in the rotated three-space. (This shows that a spatial rotation of  $\mathbf{F}$  makes a spatial rotation of  $\mathbf{E}$  and  $\mathbf{B}$ .)

c) Define the current four-vector  $\vec{J}$  by  $J^0 = \rho$ ,  $J^i = (\mathbf{J})^i$ , and show that two of the Maxwell's equations are just

$$F^{\mu\nu}_{,\nu} = 4\pi J^{\mu} .$$

d) Show that the other two of Maxwell's equations are

$$F_{\mu\nu,\lambda} + F_{\nu\lambda,\mu} + F_{\lambda\mu,\nu} = 0 .$$

Note that there are only four independent equations here. That is, choose one index value, say zero. Then the three other values (1,2,3) can be assigned to  $\mu, \nu, \lambda$  in *any* order, producing the same equation (up to an overall sign) each time. Try it and see: it follows from antisymmetry of  $F_{\mu\nu}$ .

e) We have now expressed Maxwell's equations in tensor form. Show that conservation of charge,  $J^{\mu}_{,\mu} = 0$  is implied by part c) above. (Hint: use the antisymmetry of  $F_{\mu\nu}$ .)

f) The charge density in any frame is  $J^0$ . Therefore the total charge in spacetime is  $Q = \int J^0 dx dy dz$ , where the integral extends over an entire hypersurface  $t = \text{constant}$ . Defining  $\tilde{d}t = \tilde{n}$ , a unit normal for this hypersurface, show that

$$Q = \int J^{\alpha} n_{\alpha} dx dy dz .$$

g) Use Gauss's law and the result in part c) to show that the total energy enclosed within any closed two-surface  $\mathcal{S}$  in the hypersurface  $t = \text{constant}$  can be determined by doing an integral over  $\mathcal{S}$  itself:

$$Q = \oint_{\mathcal{S}} F^{0i} n_i d\mathcal{S} = \oint \mathbf{E} \cdot \mathbf{n} d\mathcal{S} ,$$

where  $\mathbf{n}$  is the unit normal to  $\mathcal{S}$  in the hypersurface (*not* the same as  $\tilde{n}$  in part f) above).

h) Perform a Lorentz transformation on  $F^{\mu\nu}$  to a frame  $\bar{\mathcal{O}}$  moving with velocity  $v$  in the  $x$  direction relative to the frame used in part a) above. In this frame define a three-vector  $\bar{\mathbf{E}}$  with components  $\bar{E}^i = F^{\bar{0}i}$ , and similarly for  $\bar{\mathbf{B}}$  in analogy with a). In this way discover how  $\mathbf{E}$  and  $\mathbf{B}$  behave under a Lorentz transformation: they get mixed together! Thus,  $\mathbf{E}$  and  $\mathbf{B}$  themselves are not Lorentz invariant, but are merely components of  $\mathbf{F}$ , called the Faraday tensor, which is the invariant description of electromagnetic fields in relativity. If you think carefully, you will see that on physical grounds they *cannot* be invariant. In particular, the magnetic field is created by moving charges; but a charge moving in one frame may be at rest in another, so a magnetic field which exists in one frame may not exist in another. What is the same in *all* frames is the Faraday tensor: only its components get transformed.

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