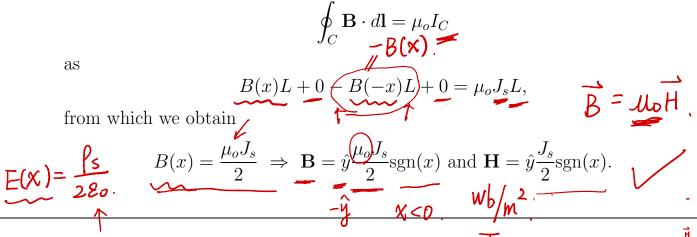
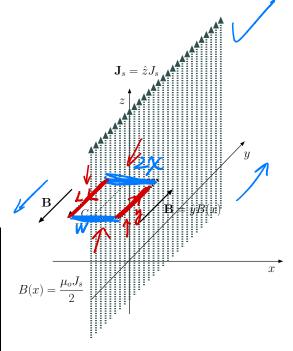
13 Current sheet, solenoid, vector potential and current loops

In the following examples we will calculate the magnetic fields $\mathbf{B} = \mu_o \mathbf{H}$ established by some simple current configurations by using the integral form of static Ampere's law. $\oint \vec{B} \cdot d\vec{\ell} = \mathcal{U}_{\sigma} \cdot \vec{L}_{\sigma}$.

Example 1: Consider a uniform surface current density $J_s = J_s \hat{z} \underline{A/m}$ flowing on x = 0 plane (see figure in the margin) — the current sheet extends infinitely in y and z directions. Determine \mathbf{B} and \mathbf{H} .

Solution: Since the current sheet extends infinitely in y and z directions we expect \mathbf{B} to depend only on coordinate x. Also, the field should be the superposition of the fields of an infinite number of current filaments, which suggests, by right-hand-rule, $\mathbf{B} = \hat{y}B(x)$, where B(x) is an odd function of x. To determine B(x), such that B(-x) = -B(x), we apply Ampere's law by computing the circulation of \mathbf{B} around the rectangular path C shown in the figure in the margin. We expand





As shown in Example 1 magnetic field of a current sheet is independent of distance |x| from the current sheet. Also H changes discontinuously across the current sheet by an amount J_s .

$$\overrightarrow{H} = \underbrace{\overset{J_{2}}{\underbrace{j}}}_{2} \widehat{j} \qquad \bigotimes \overrightarrow{H} = \underbrace{\overset{J_{3}}{\underbrace{j}}}_{2} \widehat{j} \qquad \bigotimes \underbrace{H = \underbrace{\overset{J_{5}}{\underbrace{J_{5}}}}_{2}}_{H}$$

the H field, which is tangential to the surface current density, jumps by the amount of surface current that you go through.

$$D = \frac{P_1}{2}$$

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Example 2: Consider a slab of thickness W over $-\frac{W}{2} < x < \frac{W}{2}$ which extends infinitely in y and z directions and conducts a uniform current density of $\mathbf{J} = \hat{z}J_o$ A/m^2 . Determine **H** if the current density is zero outside the slab.

Solution: Given the geometric similarities between this problem and Example 1, we postulate that $\mathbf{B} = \hat{y}B(x)$, where B(x) is an odd function of x, that is B(-x) =-B(x). To determine B(x) we apply Ampere's law by computing the circulation of $\bf B$ around the rectangular path C shown in the figure in the margin. For $x<\frac{W}{2}$, we expand

as
$$\oint_C \mathbf{B} \cdot d\mathbf{l} = \mu_o I_C$$

$$B(x)L + 0 - B(-x)L + 0 = \mu_o J_o 2xL \Rightarrow B(x) = \mu_o J_o x.$$
For $x \ge \frac{W}{2}$, the expansion gives

$$\sqrt{B(x)L + 0 - B(-x)L + 0} = \mu_o J_o WL \Rightarrow B(x) = \mu_o J_o \frac{W}{2}.$$

Hence, we find that

$$x \geq \frac{W}{2}, \text{ the expansion gives}$$

$$A = \frac{W}{2}$$

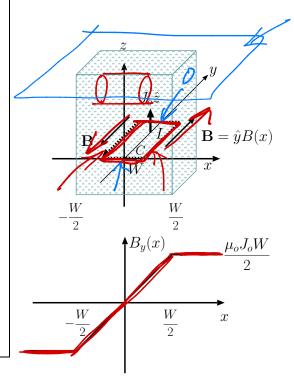
$$B(x) = \mu_o J_o \frac{W}{2}$$

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$$A = \frac{W}{2}$$

Note that the solution plotted in the margin shows no discontinuity at $x=\pm \frac{W}{2}$ or elsewhere.



The figure in the margin depicts a finite section of an infinite solenoid. A solenoid can be constructed in practice by winding a long wire into a multi loop coil as depicted. A solenoid with its loop carrying a current I in $\hat{\phi}$ direction (as shown), produces effectively a surface current density of $\mathbf{J}_s = IN\hat{\phi}$ A/m, where N is the number density (1/m) of current loops in the solenoid. In Example 3 we compute the magnetic field of the infinite solenoid using Ampere's law.

Example 3: An infinite **solenoid** having N loops per unit length is stacked in z-direction, each loop carrying a current of I A in counter-clockwise direction when viewed from the top (see margin). Determine \mathbf{H} .

Solution: Assuming that $\mathbf{B}=0$ outside the solenoid, and also \mathbf{B} is independent of z within the solenoid, we find that Ampere's law indicates for the circulation C shown in the margin

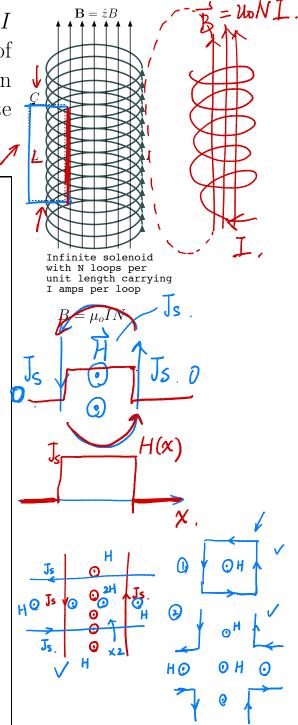
This leads to

$$B = \mu_0 IN$$
 and $\mathbf{H} = \hat{z}IN$

for the field within the solenoid.

The assumption of zero magnetic flux density $\mathbf{B} = 0$ for the exterior region is justified because:

- (a) if the exterior field is non-zero, then it must be independent of x and y (follows from Ampere's law applied to any exterior path C with $I_C = 0$), and
- (b) the finite interior flux $\Psi = \mu_o I N \pi a^2$ can only be matched with the flux of the infinitely extended exterior region when the *constant* exterior flux density (because of (a)) is vanishingly small.



• Static electric fields: Curl-free and are governed by

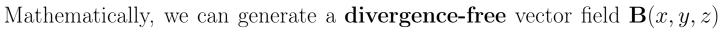
$$\nabla \times \mathbf{E} = 0, \ \nabla \cdot \mathbf{D} = \rho \text{ where } \mathbf{D} = \epsilon \mathbf{E}$$

with $\epsilon = \epsilon_r \epsilon_o$.

• Static magnetic fields: Divergence-free and are governed by

$$\nabla \cdot \mathbf{B} = 0$$
, $\nabla \times \mathbf{H} = \mathbf{J}$ where $\mathbf{B} = \mu \mathbf{H}$

with $\mu = \mu_r \mu_o$ — relative permeabilities μ_r other than unity (for free space) will be explained later on.



as

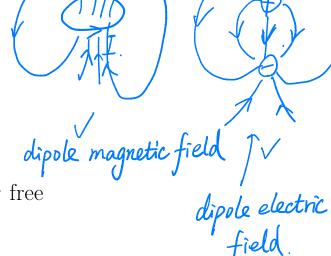
$$\nabla \cdot \left(\text{//B} = \nabla \times \mathbf{A} \right) \leftarrow$$

by taking the curl of any vector field $\mathbf{A} = \mathbf{A}(x, y, z)$ (just like we can generate a curl-free **E** by taking the gradient of any scalar field -V(x,y,z)).

Verification: Notice that

$$\nabla \cdot \nabla \times \mathbf{A} = \frac{\partial}{\partial x} (\nabla \times \mathbf{A})_x + \frac{\partial}{\partial y} (\nabla \times \mathbf{A})_y + \frac{\partial}{\partial z} (\nabla \times \mathbf{A})_z = \begin{vmatrix} \frac{\partial}{\partial x} & \frac{\partial}{\partial y} & \frac{\partial}{\partial z} \\ \frac{\partial}{\partial x} & \frac{\partial}{\partial y} & \frac{\partial}{\partial z} \\ A_x & A_y & A_z \end{vmatrix} \\
= \frac{\partial}{\partial x} (\frac{\partial A_z}{\partial y} - \frac{\partial A_y}{\partial z}) - \frac{\partial}{\partial y} (\frac{\partial A_z}{\partial x} - \frac{\partial A_x}{\partial z}) + \frac{\partial}{\partial z} (\frac{\partial A_y}{\partial x} - \frac{\partial A_x}{\partial y}) = 0.$$

• If $\mathbf{B} = \nabla \times \mathbf{A}$ represents a magnetostatic field, then \mathbf{A} is called magnetostatic potential or vector potential.



- Vector potential \mathbf{A} can be used in magnetostatics in similar ways to how electrostatic potential V is used in electrostatics.
 - In electrostatics we can assign V = 0 to any point in space that is convenient in a given problem.
 - In magnetostatics we can assign $\nabla \cdot \mathbf{A}$ to any scalar that is convenient in a given problem.
- For example, if we make the assignment¹

then we find that
$$\nabla \times \mathbf{A} = \mathbf{B}$$

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$$\nabla \times \mathbf{B} = \nabla \times \nabla \times \mathbf{A} = \nabla(\nabla \cdot \mathbf{A}) - \nabla^2 \mathbf{A} = -\nabla^2 \mathbf{A}.$$

This is a *nice and convenient* outcome, because, when combined with

$$\nabla \times \mathbf{H} = \mathbf{J} \quad \Rightarrow \quad \nabla \times \mathbf{B} = \mu_o \mathbf{J},$$

it produces

solve
$$\sqrt{\nabla^2 \mathbf{A}} = -\mu_o \mathbf{J}$$
, for \overrightarrow{A} . then $\overrightarrow{B} = \nabla \times \overrightarrow{A}$.

which is the magnetostatic version of Poisson's equation

solve
$$\nabla^2 V = -\frac{\rho}{\epsilon_o}$$
 for V . then $\vec{E} = -\nabla V$.

¹With this assignment — known as Coulomb gauge — **A** acquires the physical meaning of "potential momentum per unit charge", just as scalar potential V is "potential enegy per unit charge" (see Konopinski, $Am.\ J.\ Phys.,\ 46,\ 499,\ 1978$).

- In analogy with solution

$$V(\mathbf{r}) = \int \frac{\rho(\mathbf{r}')}{4\pi\epsilon_o|\mathbf{r} - \mathbf{r}'|} d^3\mathbf{r}' \qquad \frac{\rho}{\varepsilon_o}$$

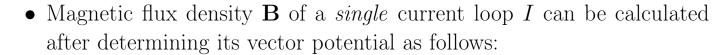
of Poisson's equation, it has a solution

$$\mathbf{A}(\mathbf{r}) = \int \frac{\mu_o \mathbf{J}(\mathbf{r}')}{4\pi |\mathbf{r} - \mathbf{r}'|} d^3\mathbf{r}'$$

Given any static² current density $\mathbf{J}(\mathbf{r})$, the above equation can be used to obtain the corresponding vector potential \mathbf{A} that simultaneously satisfies

$$\nabla \cdot \mathbf{A} = 0$$
 and $\nabla \times \mathbf{A} = \mathbf{B}$.

Once **A** is available, obtaining $\mathbf{B} = \nabla \times \mathbf{A}$ is then just a matter of taking a curl.

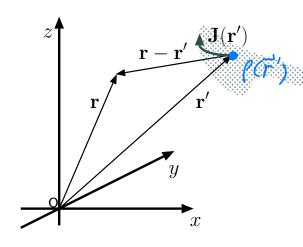


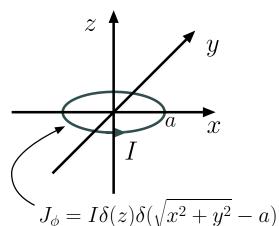
— For a loop of radius
$$a$$
 on $z=0$ plane, we can express the corresponding current density as

$$\mathbf{J}(\mathbf{r}') = I\delta(z')\delta(\sqrt{x'^2 + y'^2} - a)\frac{(-y', x', 0)}{\sqrt{x'^2 + y'^2}}$$

where the ratio on the right is the unit vector $\hat{\phi}'$.

— Inserting this into the general solution for vector potential, and performing the integration over z', we obtain





²Also, in quasi-statics we use $\mathbf{J}(\mathbf{r}',t)$ to obtain $\mathbf{A}(\mathbf{r},t)$ and $\mathbf{B} = \nabla \times \mathbf{A}$ over regions small compared to $\lambda = c/f$, with f the highest frequency in $\mathbf{J}(\mathbf{r}',t)$.

$$\mathbf{A}(\mathbf{r}) = \frac{\mu_o I}{4\pi} \int \delta(\sqrt{x'^2 + y'^2} - a) \frac{(-y', x', 0)}{\sqrt{(x - x')^2 + (y - y')^2 + z^2}} dx' dy'$$

$$= \frac{\mu_o I}{4\pi} \int \delta(r' - a) \frac{(-y', x', 0)}{\sqrt{(x - x')^2 + (y - y')^2 + z^2}} r' dr' d\phi'$$

$$= \frac{\mu_o I}{4\pi} \int_{-\pi}^{\pi} \frac{(-a \sin \phi', a \cos \phi', 0)}{\sqrt{(x - a \cos \phi')^2 + (y - a \sin \phi')^2 + z^2}} d\phi' \equiv \hat{x} A_x(\mathbf{r}) + \hat{y} A_y(\mathbf{r}).$$

- Given that $A_z = 0$, it can be shown that $\mathbf{B} = \nabla \times \mathbf{A}$ leads to

$$B_x = -\frac{\partial A_y}{\partial z}, \ B_y = \frac{\partial A_x}{\partial z}, \ B_z = \frac{\partial A_y}{\partial x} - \frac{\partial A_x}{\partial y}.$$

— From the expected azimuthal symmetry of **B** about the z-axis, it is sufficient to evaluate these on, say, y = 0 plane — after some algebra, and dropping the primes, we find, on y = 0 plane,

$$B_x = \frac{\mu_o aI}{4\pi} \int_{-\pi}^{\pi} \frac{z \cos \phi}{(x^2 + a^2 + z^2 - 2ax \cos \phi)^{3/2}} d\phi,$$

$$B_y = \frac{\mu_o aI}{4\pi} \int_{-\pi}^{\pi} \frac{z \sin \phi}{(x^2 + a^2 + z^2 - 2ax \cos \phi)^{3/2}} d\phi,$$

and

$$B_z = \frac{\mu_o aI}{4\pi} \int_{-\pi}^{\pi} \frac{a - x \cos \phi}{(x^2 + a^2 + z^2 - 2ax \cos \phi)^{3/2}} d\phi.$$

- We note that $B_y = 0$ since the B_y integrand above is odd in ϕ and the integration limits are centered about the origin. Hence, the field on y = 0 plane is given as

$$\mathbf{B} = \hat{x}B_x + \hat{z}B_z$$

with B_x and B_z defined above.

— There are no closed form expressions for the B_x and B_z integrals above for an arbitrary (x, z).

• However, it can be easily seen that if x = 0 (i.e., along the z-axis), $B_x = 0$ (as symmetry would dictate) and

$$B_z = \frac{\mu_o aI}{4\pi} \int_{-\pi}^{\pi} \frac{a}{(a^2 + z^2)^{3/2}} d\phi = \frac{\mu_o I a^2}{2(a^2 + z^2)^{3/2}}.$$

For $|z| \gg a$,

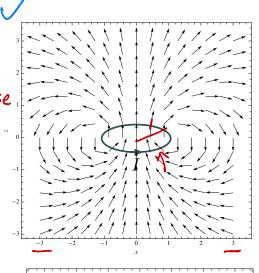
$$B_z \approx \frac{\mu_o I a^2}{2|z|^3},$$

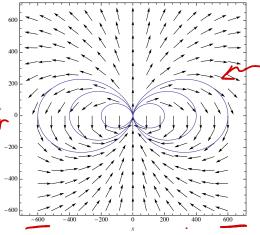
which is positive and varies with the inverse third power of distance |z|.

- Also, B_x and B_z integrals can be performed numerically. Figure in the margin depicts the pattern of \hat{B} on y=0 plane for a loop of radius a=1 computed using *Mathematica*.
- Note that circulation $\oint_C \mathbf{B} \cdot d\mathbf{l}$ around each closed field line ("linking" the current loop) equals a fixed value of $\mu_o I$ this dictates that the average field strength $|\mathbf{B}|$ of a current loop is stronger on shorter field lines closer to the current loop than on longer field lines linking the loop further out. As a result $|\mathbf{B}|$ can be shown to vary as r^{-3} for large r.
- It can be shown that the equations for magnetic *field lines* of a current loop on, say, y = 0 plane, can be expressed as

$$r = L\sin^2\theta$$

in terms of radial distance r from the origin and zenith angle θ measured from the z axis. Clearly, parameter L in this formula is the





radial distance of the field line on $\theta = 90^{\circ}$ plane, and the field line formula is accurate only for $r \gg a$. The Earth's magnetic field had such a **magnetic dipole** topology as shown.

• Lorentz force due to the magnetic fields of a pair of current loops — also known as *magnetic dipoles* — turns out to be "attractive" when the current directions agree (see margin). Bar magnets carrying "equivalent" current loops of atomic origins interact with one another in exactly the same way — i.e., as governed by the second term of Lorentz force.

