

## Lecture 2: Random Signals

Mark Hasegawa-Johnson

ECE 537: Speech Processing

- 1 The Problem of Phase
- 2 Power Spectrum
- 3 Autocorrelation
- 4 Cross-correlation and cross-power spectrum
- 5 Filtering
- 6 Examples
- 7 Summary

# Outline

- 1 The Problem of Phase
- 2 Power Spectrum
- 3 Autocorrelation
- 4 Cross-correlation and cross-power spectrum
- 5 Filtering
- 6 Examples
- 7 Summary

# What time is zero?

The DTFT is defined to be

$$X(\omega) = \sum_{n=-\infty}^{\infty} x[n]e^{-j\omega n}$$

But what time is zero?

- Buddhist zero: Midnight 544BCE, Kushinagar
- Christian zero: Midnight 0CE, Bethlehem
- Unix: Midnight UTC, January 1, 1970
- Most students: whatever time I started the recording

... does it matter?

# Unfortunately, it matters

Suppose  $y[n] = x[n - n_0]$ , i.e., they are the same signal, but measured with different definitions of  $n = 0$ .

$$Y(\omega) = X(\omega)e^{-j\omega n_0}$$

$$Y[k] = X[k]e^{-j\frac{2\pi kn_0}{N}}$$

- If  $\left|\frac{2\pi n_0}{N}\right| < \pi$ , the starting time of your recording can be recovered from data using **phase unwrapping**:

$$n_0 = \underset{n_0}{\operatorname{argmin}} \left| \angle Y[k] - \left( \angle Y[k-1] - \frac{2\pi n_0}{N} \right) \right|$$

- If  $n_0 > 0.5N$ , phase unwrapping finds, at best, an aliased start time. In that case, phase is very hard to model – random, for most practical purposes.

# How do we deal with phase?

There are two standard methods for dealing with phase:

- **Metadata:** Keep careful metadata specifying how  $n = 0$  is defined. The most common example: Pitch-synchronous analysis (e.g., PSOLA=pitch-synchronous overlap-and-add) defines  $n = 0$  at the moment of a glottal closure event.
- **Phase-invariant features:** Ignore  $\angle X(\omega)$ . Instead, work with only  $|X(\omega)|$ , which is invariant to phase.

# Outline

- 1 The Problem of Phase
- 2 Power Spectrum**
- 3 Autocorrelation
- 4 Cross-correlation and cross-power spectrum
- 5 Filtering
- 6 Examples
- 7 Summary

# Magnitude DTFT is time-invariant, magnitude DFT is not

The magnitude DTFT is time-invariant. If  $y[n] = x[n - n_0]$ , then

$$|Y(\omega)| = \left| \sum_{n=-\infty}^{\infty} x[n - n_0] e^{-j\omega n} \right| = \left| e^{-j\omega n_0} \sum_{\ell=-\infty}^{\infty} x[\ell] e^{-j\omega \ell} \right|$$

The magnitude DFT is NOT time-invariant:

$$|Y[k]| = \left| \sum_{n=0}^{N-1} x[n - n_0] e^{-j\frac{2\pi kn}{N}} \right| \neq \left| e^{-j\frac{2\pi kn_0}{N}} \sum_{\ell=0}^{N-1} x[\ell] e^{-j\frac{2\pi k\ell}{N}} \right|$$

By windowing the signal, we have selected  $N$  random samples.

$|Y[k]|$  is therefore a random variable; it has the same distribution as  $|X[k]|$ , but not the same value.

# Power Spectrum

Power spectrum is defined as the expected squared magnitude of the Fourier transform:

$$P_{XX}(\omega) = E \left[ |X(\omega)|^2 \right]$$

$$P_{XX}[k] = E \left[ |X[k]|^2 \right]$$

Finally! The power spectrum is a time-shift-invariant feature.

# Power Spectral Estimators

Unfortunately, the power spectrum can't be computed: expectation ( $E[\cdot]$ ) is the average over infinitely many instances of the random signal. There are three most common estimates:

- Periodogram (noisy but easy):

$$\hat{P}_{XX}[k] = \left| \sum_{n=0}^{N-1} x[n] e^{-j \frac{2\pi kn}{N}} \right|^2$$

- Bartlett's Method (average of the periodograms of multiple frames):

$$\hat{P}_{XX}[k] = \frac{1}{T} \sum_{t=0}^{T-1} \left| \sum_{n=0}^{N-1} x[n - tN] e^{-j \frac{2\pi kn}{N}} \right|^2$$

- Welch's Method = average of the periodograms of windowed overlapping frames.

# Outline

- 1 The Problem of Phase
- 2 Power Spectrum
- 3 Autocorrelation**
- 4 Cross-correlation and cross-power spectrum
- 5 Filtering
- 6 Examples
- 7 Summary

# Expectation is Linear

The expectation operator is linear. Remember that if  $p_{X,Y}(x,y)$  is the joint pdf of random variables  $X$  and  $Y$ , then

$$\begin{aligned} E[aX + bY] &= \int \int (ax + by) p_{X,Y}(x,y) dx dy \\ &= aE[X] + bE[Y] \end{aligned}$$

# Periodogram is DFT of a Correlation

The periodogram is the magnitude squared DFT, which are samples of the magnitude-squared DTFT:

$$\hat{P}_{XX}(\omega) = |X(\omega)|^2 = X(\omega)X^*(\omega)$$

Multiplication and convolution are a Fourier transform pair:

$$X(\omega)X^*(\omega) \leftrightarrow x[n] * z[n],$$

where  $z[n] = x^*[-n]$ :

$$\begin{aligned} z[n] &= \frac{1}{2\pi} \int_0^{2\pi} X^*(\omega) e^{j\omega n} \\ &= \left( \frac{1}{2\pi} \int_0^{2\pi} X(\omega) e^{j\omega(-n)} \right)^* \\ &= (x[-n])^* \end{aligned}$$

# Power Spectrum $\leftrightarrow$ Autocorrelation

Expectation is linear. Fourier transform is linear. Convolution is the sum of products. So we can swap their order:

$$\begin{aligned} P_{XX}(\omega) &= E [|X(\omega)|^2] \\ &= E [\text{DTFT} \{x[n] * x^*[-n]\}] \\ &= E \left[ \text{DTFT} \left\{ \sum_{n=-\infty}^{\infty} x[n] x^*[n-m] \right\} \right] \\ &= \text{DTFT} \left\{ \sum_{n=-\infty}^{\infty} E [x[n] x^*[n-m]] \right\} \\ &= \text{DTFT} \{ E [x[n] x^*[n-m]] \} \end{aligned}$$

where the last line assumes that the signal is **wide-sense stationary** (WSS), meaning that its autocorrelation is independent of  $n$ .

# The power spectrum is the DTFT of the autocorrelation

$$P_{XX}(\omega) \equiv E [|X(\omega)|^2]$$

$$R_{xx}[m] \equiv E [x[n]x^*[n - m]]$$

$$P_{XX}(\omega) = \text{DTFT} \{R_{XX}[m]\}$$

# Properties of the autocorrelation

- **Conjugate symmetric:**

$$\begin{aligned} R_{XX}[-m] &= E [x[n]x^*[n+m]] \\ &= E [x[\ell-m]x^*[\ell]], \quad \ell = n+m \\ &= R_{XX}^*[m] \end{aligned}$$

If  $x[n]$  is real, then  $R_{XX}[m]$  is real and symmetric.

- $R_{XX}[0]$  is the **signal power**:

$$\begin{aligned} R_{XX}[0] &= E [x[n]x^*[n]] \\ &= E [|x[n]|^2] \end{aligned}$$

# Autocorrelation of a Sinusoid

As a first example, let's find the autocorrelation and power spectrum of a sinusoid. Suppose  $x[n] = A \cos(\omega_0 n + \theta)$ . Suppose  $\omega_0$  and  $A$  are known, and only  $\theta$  is unknown (because we don't know how  $n = 0$  is defined).

$$\begin{aligned} R_{XX}[m] &= E[x[n]x^*[n-m]] \\ &= A^2 E[\cos(\omega_0 n + \theta) \cos(\omega_0(n-m) + \theta)] \end{aligned}$$

Using the trig formula  $\cos(a) \cos(b) = \frac{1}{2} (\cos(a+b) + \cos(a-b))$ , we get

$$R_{XX}[m] = \frac{A^2}{2} E[\cos(\omega_0(2n-m) + 2\theta) + \cos(\omega_0 m)]$$

But if  $\theta$  is uniformly distributed, then  $E[\cos \theta] = 0$ , so

$$R_{XX}[m] = \frac{A^2}{2} \cos(\omega_0 m)$$

Notice that's real and symmetric!

# Power Spectrum a Sinusoid

The power spectrum of  $x[n] = A \cos(\omega_0 n + \theta)$  is

$$\begin{aligned} P_{XX}(\omega) &= \text{DTFT} \left\{ \frac{A^2}{2} \cos(\omega_0 m) \right\} \\ &= \frac{\pi A^2}{2} (\delta(\omega - \omega_0) + \delta(\omega + \omega_0)) \end{aligned}$$

# Autocorrelation of White Noise

As a second example, let's find the autocorrelation of white noise. "White noise" is a type of signal whose power spectrum is a constant, independent of  $\omega$ , like white light:

$$P_{XX}(\omega) = 1$$

The autocorrelation of such a signal is the inverse transform, which is a delta function:

$$\begin{aligned} R_{XX}[m] &= \delta[m] \\ &= \begin{cases} 1 & m = 0 \\ 0 & m \neq 0 \end{cases} \\ &= E[x[n]x^*[n-m]] \end{aligned}$$

In words: A white noise signal is one with uncorrelated samples.

# Outline

- 1 The Problem of Phase
- 2 Power Spectrum
- 3 Autocorrelation
- 4 Cross-correlation and cross-power spectrum
- 5 Filtering
- 6 Examples
- 7 Summary

# Cross-Correlation

Suppose we have two signals,  $x[n]$  and  $y[n]$ , that are both random, not equal to each other, but also not completely uncorrelated. We can write their cross-correlation as

$$R_{XY}[m] = E [x[n]y^*[n - m]]$$

The cross-power spectrum is

$$\begin{aligned} P_{XY}(\omega) &= \text{DTFT} \{R_{XY}[m]\} \\ &= E [\text{DTFT} \{x[n] * y^*[-n]\}] \\ &= E [X(\omega)Y^*(\omega)] \end{aligned}$$

# Properties of the Cross-Correlation and Cross-power spectrum

- **Cross-correlation: conjugate antisymmetric in RVs and time:**

$$\begin{aligned} R_{XY}[m] &= E [x[n]y^*[n - m]] \\ &= (E [y[\ell]x^*[\ell + m]])^* , \quad \ell = n - m \\ &= R_{YX}^*[-m] \end{aligned}$$

- **Cross-power spectrum: conjugate antisymmetric in RVs:**

$$\begin{aligned} P_{XY}(\omega) &= E [X(\omega)Y^*(\omega)] \\ &= P_{YX}^*(\omega) \end{aligned}$$

# Cross-power spectra of two sinusoids at the same frequency

Suppose, first, that  $x[n] = A \cos(\omega_0 n + \theta)$  and  $y[n] = B \cos(\omega_0 n + \phi)$ .

$$\begin{aligned} R_{XY}[m] &= AB \times E [\cos(\omega_0 n + \theta) \cos(\omega_0(n - m) + \phi)] \\ &= \frac{AB}{2} \cos(\omega_0 m + (\theta - \phi)) \end{aligned}$$

The cross-power spectrum is

$$P_{XY}(\omega) = \frac{AB\pi}{2} \left( e^{j(\theta-\phi)} \delta(\omega - \omega_0) + e^{-j(\theta-\phi)} \delta(\omega + \omega_0) \right)$$

# Cross-power spectra of signal and shifted signal

Generalizing the previous slide: suppose  $y[n] = ax[n - n_0] + v[n]$ , where  $v[n]$  is some noise signal that is uncorrelated with  $x[n]$ . Then

$$R_{XY}[m] = E [x[n](ax^*[n - n_0 - m] + v^*[n - m])]$$

$$\begin{aligned} R_{XY}[m] &= aE [x[n]x^*[n - n_0 - m]] + E [x[n]v^*[n - m]] \\ &= aR_{XX}[m + n_0] + 0 \end{aligned}$$

The cross-power spectrum is

$$P_{XY}(\omega) = ae^{j\omega n_0} P_{XX}(\omega)$$

# Outline

- 1 The Problem of Phase
- 2 Power Spectrum
- 3 Autocorrelation
- 4 Cross-correlation and cross-power spectrum
- 5 Filtering**
- 6 Examples
- 7 Summary

# Real-world signals

- In the real world, pure tones and pure white-noise signals are rare.
- Most real-world signals can be modeled as  $y[n] = h[n] * x[n]$ , where  $x[n]$  is a sum of pure tones and/or white noise, and  $h[n]$  is a spectral shaping filter.
- Let's derive the cross-correlation, autocorrelation, cross-power spectrum, and power spectrum.

# Cross-correlation

Suppose  $y[n] = h[n] * x[n]$ , then

$$\begin{aligned} R_{YX}[m] &= E[y[n]x^*[n-m]] \\ &= E\left[\left(\sum_{\ell=-\infty}^{\infty} h[\ell]x[n-\ell]\right)x^*[n-m]\right] \\ &= \sum_{\ell=-\infty}^{\infty} h[\ell]E[x[n-\ell]x^*[n-m]] \\ &= \sum_{\ell=-\infty}^{\infty} h[\ell]R_{XX}[m-\ell] \\ &= h[m] * R_{XX}[m] \end{aligned}$$

$$R_{XY}[m] = R_{YX}^*[-m] = h^*[-m] * R_{XX}[m]$$

# Cross-power spectrum

Suppose  $y[n] = h[n] * x[n]$ , then

$$R_{XY}[m] = h^*[-m] * R_{XX}[m]$$

$$R_{YX}[m] = h[m] * R_{XX}[m]$$

Taking the DTFT, we get:

$$P_{XY}(\omega) = H^*(\omega)P_{XX}(\omega)$$

$$P_{YX}(\omega) = H(\omega)P_{XX}(\omega)$$

# Autocorrelation

Suppose  $y[n] = h[n] * x[n]$ , then

$$\begin{aligned} R_{YY}[m] &= E[y[n]y^*[n-m]] \\ &= E\left[\left(\sum_{\ell=-\infty}^{\infty} h[\ell]x[n-\ell]\right)y^*[n-m]\right] \\ &= \sum_{\ell=-\infty}^{\infty} h[\ell]E[x[n-\ell]y^*[n-m]] \\ &= \sum_{\ell=-\infty}^{\infty} h[\ell]R_{XY}[m-\ell] \\ &= h[m] * R_{XY}[m] \\ &= h[m] * h^*[-m] * R_{XX}[m] \end{aligned}$$

The time-reversed convolution ( $h[m] * h^*[-m]$ ) is conjugate antisymmetric.

# Power Spectrum

Suppose  $y[n] = h[n] * x[n]$ , then

$$\begin{aligned} P_{YY}(\omega) &= \text{DTFT} \{R_{YY}[m]\} \\ &= \text{DTFT} \{h[m] * h^*[-m] * R_{XX}[m]\} \\ &= H(\omega)H^*(\omega)P_{XX}(\omega) \\ &= |H(\omega)|^2 P_{XX}(\omega) \end{aligned}$$

# Outline

- 1 The Problem of Phase
- 2 Power Spectrum
- 3 Autocorrelation
- 4 Cross-correlation and cross-power spectrum
- 5 Filtering
- 6 Examples**
- 7 Summary

# Examples

- 1 Suppose  $x[n]$  is unit-variance white noise, and  $y[n] = x[n] + ax[n - 1]$ . What is  $P_{YY}(\omega)$ ?
- 2 Suppose  $x[n]$  is unit-variance white noise, and  $y[n] = x[n] - ay[n - 1]$ . What is  $P_{YY}(\omega)$ ?

# Outline

- 1 The Problem of Phase
- 2 Power Spectrum
- 3 Autocorrelation
- 4 Cross-correlation and cross-power spectrum
- 5 Filtering
- 6 Examples
- 7 Summary**

# Summary

The power spectrum of white noise is constant; the power spectrum of a sinusoid is a cosine. If  $y[n] = h[n] * x[n]$ , then

$$R_{YX}[m] = h[m] * R_{XX}[m]$$

$$R_{XY}[m] = h^*[-m] * R_{XX}[m]$$

$$R_{YY}[m] = h[m] * h^*[-m] * R_{XX}[m]$$

$$P_{YX}(\omega) = H(\omega)P_{XX}(\omega)$$

$$P_{XY}(\omega) = H^*(\omega)P_{XX}(\omega)$$

$$P_{YY}(\omega) = |H(\omega)|^2 P_{XX}(\omega)$$